

SOLUTIONS TO ASSIGNMENTS 04

1. TENSOR ANALYSIS II: THE COVARIANT DIVERGENCE AND THE LAPLACIAN

(a) The covariant divergence is $\nabla_\mu V^\mu = \partial_\mu V^\mu + \Gamma_{\mu\lambda}^\mu V^\lambda$ where

$$\Gamma_{\mu\lambda}^\mu = \frac{1}{2}g^{\mu\rho}(\partial_\mu g_{\rho\lambda} + \partial_\lambda g_{\mu\rho} - \partial_\rho g_{\mu\lambda}) = \frac{1}{2}g^{\mu\rho}\partial_\lambda g_{\mu\rho} \quad (1)$$

(the 1st and 3rd term cancel). Now we use $g^{-1}\partial_\lambda g = g^{\mu\nu}\partial_\lambda g_{\mu\nu}$ to find

$$\Gamma_{\mu\lambda}^\mu = \frac{1}{2}g^{\mu\rho}\partial_\lambda g_{\mu\rho} = \frac{1}{2}g^{-1}\partial_\lambda g = g^{-1/2}\partial_\lambda g^{+1/2} \quad (2)$$

where in the last equality we used the fact that $\partial_\lambda g^{+1/2} = \frac{1}{2}g^{-1/2}\partial_\lambda g$. We can now compute the covariant divergence

$$\nabla_\mu J^\mu = \partial_\mu J^\mu + \Gamma_{\mu\lambda}^\mu J^\lambda = \partial_\mu J^\mu + J^\lambda g^{-1/2}\partial_\lambda g^{+1/2} = g^{-1/2}\partial_\mu(g^{1/2}J^\mu) \quad (3)$$

(b) Analogously, for the covariant divergence of an anti-symmetric (2,0)-tensor $F^{\mu\nu}$ one has, using (3),

$$\begin{aligned} \nabla_\mu F^{\mu\nu} &= \partial_\mu F^{\mu\nu} + \Gamma_{\mu\rho}^\mu F^{\rho\nu} + \Gamma_{\mu\rho}^\nu F^{\mu\rho} = \partial_\mu F^{\mu\nu} + \Gamma_{\mu\rho}^\mu F^{\rho\nu} \\ &= \partial_\mu F^{\mu\nu} + F^{\rho\nu}g^{-1/2}\partial_\rho g^{+1/2} = g^{-1/2}\partial_\mu(g^{1/2}F^{\mu\nu}) \end{aligned} \quad (4)$$

where the last term in the first equation vanishes because an antisymmetric tensor ($F^{\mu\rho}$) is contracted with a symmetric object ($\Gamma_{\mu\rho}^\nu$).

Finally, using the result (3) and $\nabla_\beta f = \partial_\beta f$ the Laplacian can be written as

$$\square f = \nabla_\alpha(g^{\alpha\beta}\nabla_\beta f) = g^{-1/2}\partial_\alpha(g^{1/2}g^{\alpha\beta}\partial_\beta f) \quad (5)$$

2. STATIC SCHWARZSCHILD OBSERVERS IN EDDINGTON-FINKELSTEIN COORDINATES

In ingoing EF coordinates, the Schwarzschild metric takes the form

$$ds^2 = -f(r)dv^2 + 2dvdr + r^2d\Omega^2 \quad (6)$$

where $v = t + r_*$ satisfies

$$dv = dt + dr_* = dt + dr/f(r) \quad \Rightarrow \quad \frac{\partial v}{\partial t} = 1 \quad , \quad \frac{\partial v}{\partial r} = f(r)^{-1} \quad (7)$$

A static observer has 4-velocity

$$(u^\alpha)_{EF} = (u^v, u^r, u^\theta, u^\phi) = (u^v, 0, 0, 0) \quad , \quad (8)$$

with

$$g_{\alpha\beta}u^\alpha u^\beta = -f(r)(u^v)^2 = -1 \quad \Rightarrow \quad u^v = f(r)^{-1/2} \quad (9)$$

Thus

$$(u^\alpha)_{EF} = (f(r)^{-1/2}, 0, 0, 0) . \quad (10)$$

This agrees with the result in SS coordinates, as could have also been deduced from the vectorial transformation behaviour (y^μ are now the SS coordinates)

$$(u^v)_{EF} = \frac{\partial v}{\partial y^\mu} (u^\mu)_{SS} = \frac{\partial v}{\partial t} (u^t)_{SS} = (u^t)_{SS} . \quad (11)$$

(a) Since $\dot{u}^\alpha = 0$ for a static oberver, the acceleration is

$$a^\alpha = \Gamma_{\beta\gamma}^\alpha u^\beta u^\gamma = \Gamma_{vv}^\alpha (u^v)^2 = f(r)^{-1} \Gamma_{vv}^\alpha . \quad (12)$$

Even though Christoffel symbols are non-tensorial in general, they do transform in a simple way under the very simple coordinate transformation between SS and EF coordinates. Nevertheless it is more convenient (and in any case a good exercise) to just calculate the relevant Christoffel symbols directly in EF coordinates rather than transforming them from the result in SS coordinates.

First of all, the Christoffel symbol $\Gamma_{\alpha vv}$ is rather obviously non-zero only for $\alpha = r$, and

$$\Gamma_{rvv} = -\frac{1}{2} g_{vv,r} = +\frac{1}{2} f'(r) = m/r^2 . \quad (13)$$

To determine Γ_{vv}^α we need the components of the inverse metric. Because the metric is not diagonal, this requires a bit of care. In matrix form, the (v, r) -components of the metric and its inverse are

$$(g_{\alpha\beta}) = \begin{pmatrix} -f & 1 \\ 1 & 0 \end{pmatrix} , \quad (g^{\alpha\beta}) = \begin{pmatrix} 0 & 1 \\ 1 & f \end{pmatrix} \quad (14)$$

Therefore, there are two non-vanishing Christoffel symbols Γ_{vv}^α , namely

$$\Gamma_{vv}^v = \Gamma_{rvv} , \quad \Gamma_{vv}^r = f(r) \Gamma_{rvv} , \quad (15)$$

and the acceleration vector is

$$(a^\alpha)_{EF} = (a^v = f(r)^{-1} m/r^2, a^r = m/r^2, 0, 0) . \quad (16)$$

Thus the (non-singular, ‘‘Newtonian’’) r -component agrees with that of the acceleration in SS coordinates, but in addition in EF coordinates there is a v -component which is singular as $r \rightarrow 2m$.

(b) Alternatively, and more quickly, since a^α is a vector, one can obtain this result by transforming the result in SS coordinates to EF coordinates,

$$(a^\alpha)_{EF} = \frac{\partial x^\alpha}{\partial y^\mu} (a^\mu)_{SS} = \frac{\partial x^\alpha}{\partial r} (a^r)_{SS} \Rightarrow \begin{cases} (a^v)_{EF} = f(r)^{-1} m/r^2 \\ (a^r)_{EF} = m/r^2 \end{cases} \quad (17)$$

(c) Finally, the norm of the acceleration in EF coordinates is

$$g_{\alpha\beta} a^\alpha a^\beta = -f(r) (a^v)^2 + 2a^v a^r = f(r)^{-1} (m/r^2)^2 , \quad (18)$$

in complete agreement with the result in SS coordinates (as it should).